



Roman

1
00:00:05,269 --> 00:00:03,350
astronomy is largely the study of light

2
00:00:07,190 --> 00:00:05,279
from distant places

3
00:00:09,589 --> 00:00:07,200
each kind of light carries different

4
00:00:12,230 --> 00:00:09,599
information that adds to our picture of

5
00:00:14,310 --> 00:00:12,240
the universe

6
00:00:16,630 --> 00:00:14,320
the upcoming nancy grace roman space

7
00:00:18,950 --> 00:00:16,640
telescope will detect a key range of

8
00:00:22,230 --> 00:00:18,960
light for studying the universe

9
00:00:24,710 --> 00:00:22,240
visible into near infrared

10
00:00:28,070 --> 00:00:24,720
infrared which starts at a wavelength of

11
00:00:30,070 --> 00:00:28,080
about 0.75 microns will allow the roman

12
00:00:31,990 --> 00:00:30,080
space telescope to make many critical

13
00:00:34,229 --> 00:00:32,000

observations

14

00:00:38,150 --> 00:00:34,239

roman engineers now plan to add a new

15

00:00:39,430 --> 00:00:38,160

filter extending its range from 2 to 2.3

16

00:00:41,910 --> 00:00:39,440

microns

17

00:00:44,069 --> 00:00:41,920

this seemingly small change will make a

18

00:00:47,029 --> 00:00:44,079

big difference

19

00:00:49,510 --> 00:00:47,039

our galaxy the milky way is filled with

20

00:00:52,069 --> 00:00:49,520

bands of dust and gas that block our

21

00:00:53,830 --> 00:00:52,079

view of the stars behind them

22

00:00:56,470 --> 00:00:53,840

part of what makes infrared light so

23

00:00:58,950 --> 00:00:56,480

useful to astronomers is its ability to

24

00:01:00,950 --> 00:00:58,960

travel through this gas and dust

25

00:01:02,310 --> 00:01:00,960

infrared light has a longer wavelength

26

00:01:03,910 --> 00:01:02,320

than visible light

27

00:01:05,910 --> 00:01:03,920

which means it is less likely to be

28

00:01:07,750 --> 00:01:05,920

scattered and absorbed by small dust

29

00:01:09,510 --> 00:01:07,760

particles as it travels over long

30

00:01:12,950 --> 00:01:09,520

distances

31

00:01:15,270 --> 00:01:12,960

upgrading from 2 to 2.3 microns

32

00:01:17,990 --> 00:01:15,280

allows astronomers to see through 2 to 3

33

00:01:20,310 --> 00:01:18,000

times as much dust this opens up much

34

00:01:22,950 --> 00:01:20,320

more of our galaxy to study including

35

00:01:25,670 --> 00:01:22,960

surveys of small dim stars that glow

36

00:01:27,910 --> 00:01:25,680

mostly in infrared

37

00:01:30,550 --> 00:01:27,920

infrared is also good for studying the

38

00:01:32,950 --> 00:01:30,560

more distant parts of the universe

39

00:01:35,190 --> 00:01:32,960

as the universe expands it stretches the

40

00:01:37,190 --> 00:01:35,200

wavelengths of light along with it a

41

00:01:39,590 --> 00:01:37,200

process called red shift

42

00:01:41,190 --> 00:01:39,600

the farther away an object is the more

43

00:01:43,030 --> 00:01:41,200

the light from it has stretched by the

44

00:01:44,710 --> 00:01:43,040

time it reaches us

45

00:01:48,069 --> 00:01:44,720

distant galaxies have all of their

46

00:01:50,230 --> 00:01:48,079

visible light shifted into infrared

47

00:01:52,830 --> 00:01:50,240

stretching over distance makes redshift

48

00:01:55,190 --> 00:01:52,840

one of the key tools for measuring the

49

00:01:56,950 --> 00:01:55,200

universe since astronomers can usually

50

00:01:58,550 --> 00:01:56,960

determine what wavelength they would see

51

00:02:01,270 --> 00:01:58,560

from up close

52

00:02:04,149 --> 00:02:01,280

they can tell how far a galaxy is by how

53

00:02:06,069 --> 00:02:04,159

much the light has changed

54

00:02:08,710 --> 00:02:06,079

closer to home is the search for water

55

00:02:10,869 --> 00:02:08,720

within the solar system water ice

56

00:02:12,949 --> 00:02:10,879

absorbs specific wavelengths of infrared

57

00:02:14,390 --> 00:02:12,959

light providing a fingerprint of its

58

00:02:16,229 --> 00:02:14,400

presence

59

00:02:18,309 --> 00:02:16,239

as telescopes see farther into the

60

00:02:19,510 --> 00:02:18,319

infrared they can see more of this

61

00:02:21,589 --> 00:02:19,520

fingerprint

62

00:02:23,430 --> 00:02:21,599

if objects in the outer solar system

63

00:02:25,990 --> 00:02:23,440

such as rocky fragments in the distant

64

00:02:28,070 --> 00:02:26,000

kuiper belt contain water ice light

65

00:02:30,550 --> 00:02:28,080

reflected off them will have gaps where

66

00:02:32,869 --> 00:02:30,560

the water has absorbed that wavelength

67

00:02:36,070 --> 00:02:32,879

this allows astronomers to detect water

68

00:02:38,710 --> 00:02:36,080

at much greater distances

69

00:02:40,949 --> 00:02:38,720

the nancy grace roman space telescope

70

00:02:42,869 --> 00:02:40,959

will form an unprecedented partnership

71

00:02:44,390 --> 00:02:42,879

with hubble and the james webb space

72

00:02:47,030 --> 00:02:44,400

telescope

73

00:02:49,190 --> 00:02:47,040

with its extremely large field of view

74

00:02:51,589 --> 00:02:49,200

roman is uniquely equipped for large

75

00:02:53,830 --> 00:02:51,599

surveys of the infrared sky allowing

76

00:02:56,390 --> 00:02:53,840

astronomers to identify interesting

77

00:02:58,390 --> 00:02:56,400

targets for more detailed study using

78

00:03:01,670 --> 00:02:58,400

hubble and webb's overlapping wavelength

79

00:03:04,229 --> 00:03:01,680

ranges and smaller fields of view

80

00:03:07,730 --> 00:03:04,239

this collaboration will usher in a new